

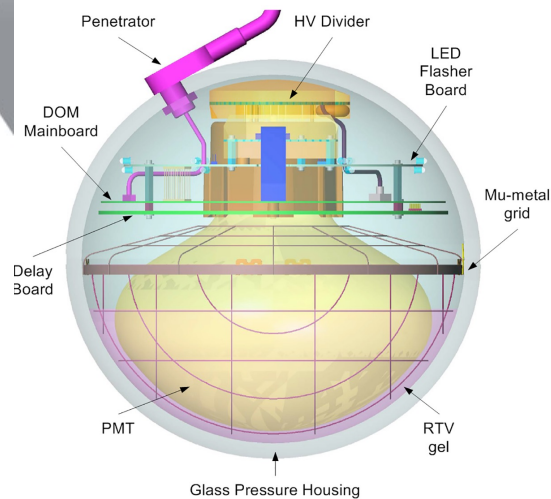
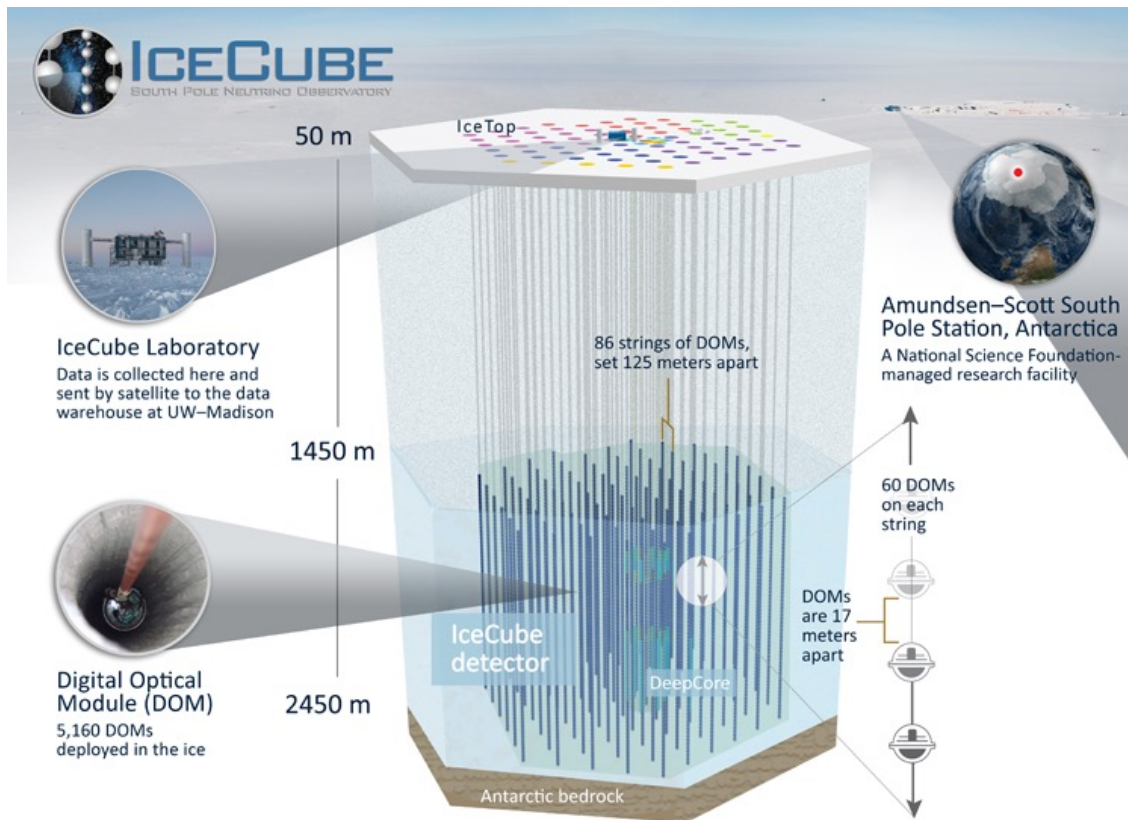
# Simulating light in scattering dominated media

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GABRIEL COLLIN

# IceCube

Gigaton neutrino detector located at the south pole



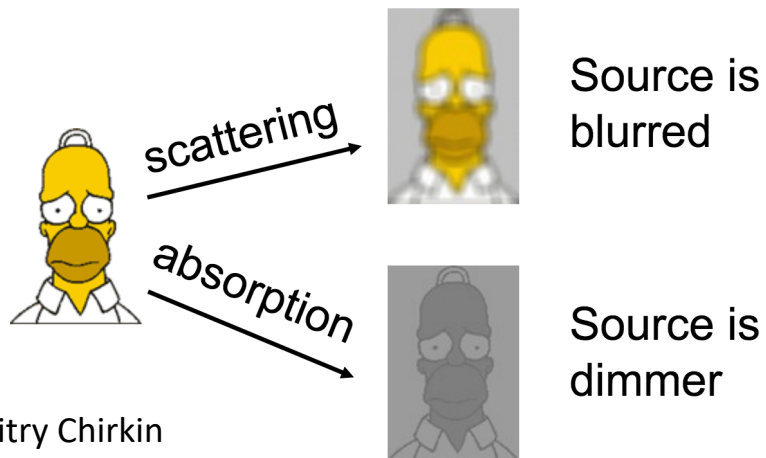
# IceCube

Muon neutrinos interact with the surrounding ice/rock and produce muons that travel through the detector.

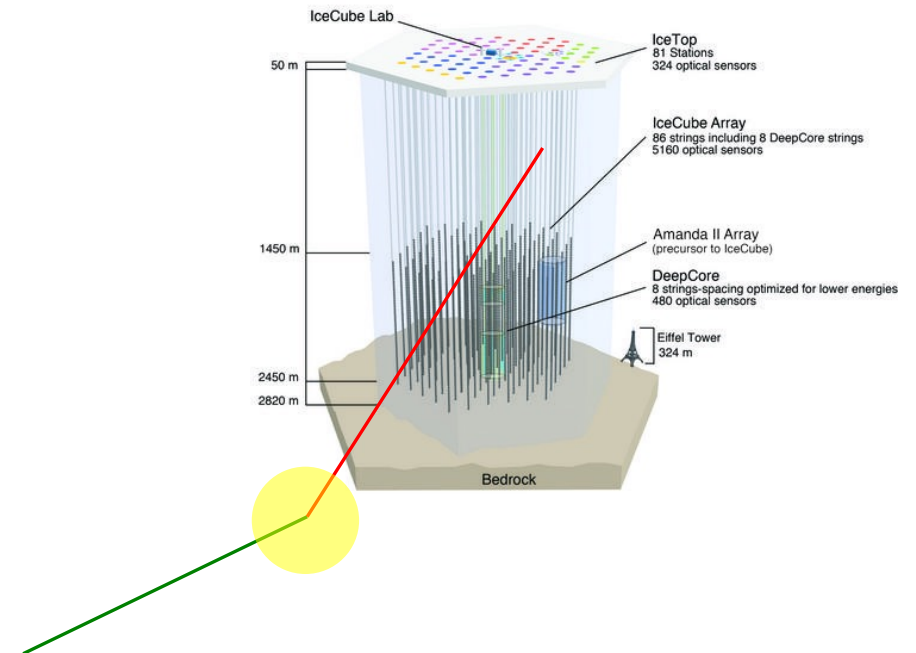
- Produces Cherenkov light as it travels.

Cherenkov light is scattered and absorbed

- Effects the angular and energy resolution.



Example by Dmitry Chirkin

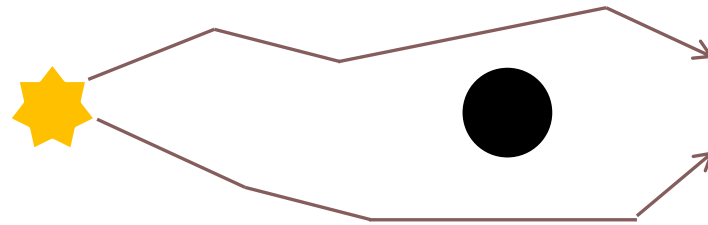


# Motivation

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IceCube uses path tracing to propagate light in the ice.

- However, most rays never reach a DOM.



Path tracers can be run *backwards* in time, but then most rays will never reach a light source.

- Path tracers can't constrain both the starting and ending location of the rays.

The fundamental problem is that the interesting paths are highly constrained.

- Is there another way to approach this?

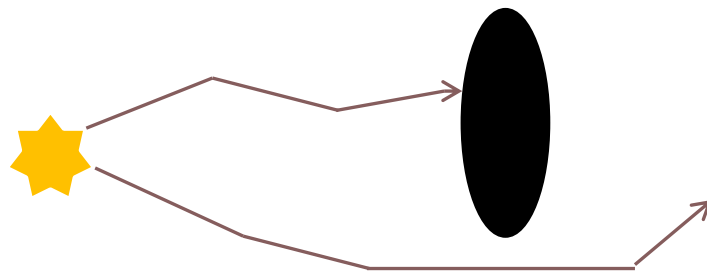
# Approximations

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## DETECTOR BIAS

Make the detectors much larger to capture more rays.

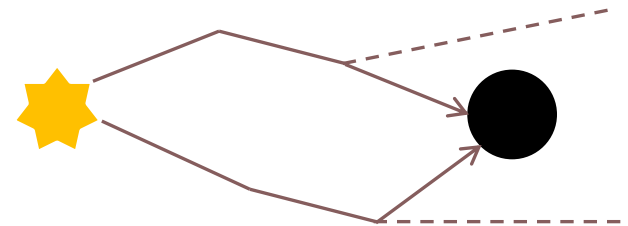
- Introduces bias that can't be corrected for.



## PATH BIAS

Adjust the path the light takes to bias it towards the detector.

- Introduces bias that can be corrected for.
- However, can massively increase variance to the point of being useless.



Can we do this without bias?

# Path integration

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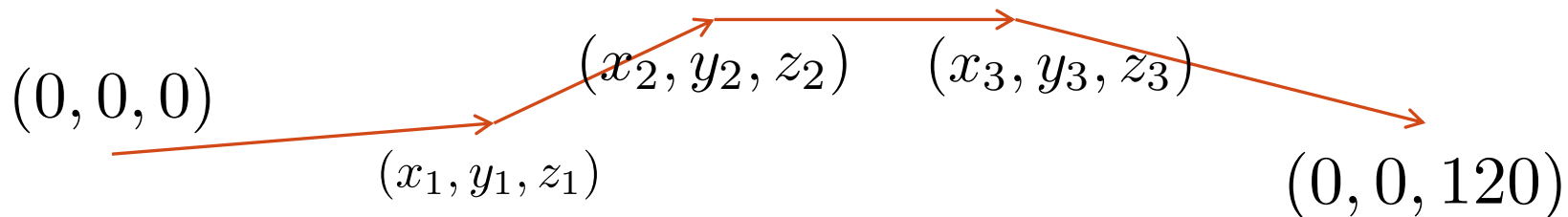
The start and end locations of the ray can be constrained if the problem is specified in terms of a classical path integral.

$$\int_{f \in \Omega} e^{-S[f]} \mathcal{D}f$$

Space of all paths

Probability of path 'f'

Eg:  $f = \{(0, 0, 0), (x_1, y_1, z_1), (x_2, y_2, z_2), \dots, (0, 0, 120)\}$



# Evaluation of the integral

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Information can be extracted about the light propagation by framing the integrand as a probability distribution:

$$e^{-S[f]} \rightarrow p[f] = p(x_1, y_1, z_1, x_2, y_2, z_2, \dots)$$

This distribution can be sampled with an MCMC.

More details on the construction of  $p[f]$  in the paper:

[arXiv.org](#) > [hep-ex](#) > [arXiv:1811.04156](#)

High Energy Physics – Experiment

**Using path integrals for the propagation of light in a scattering dominated medium**

[Gabriel H. Collin](#)

# Industry use

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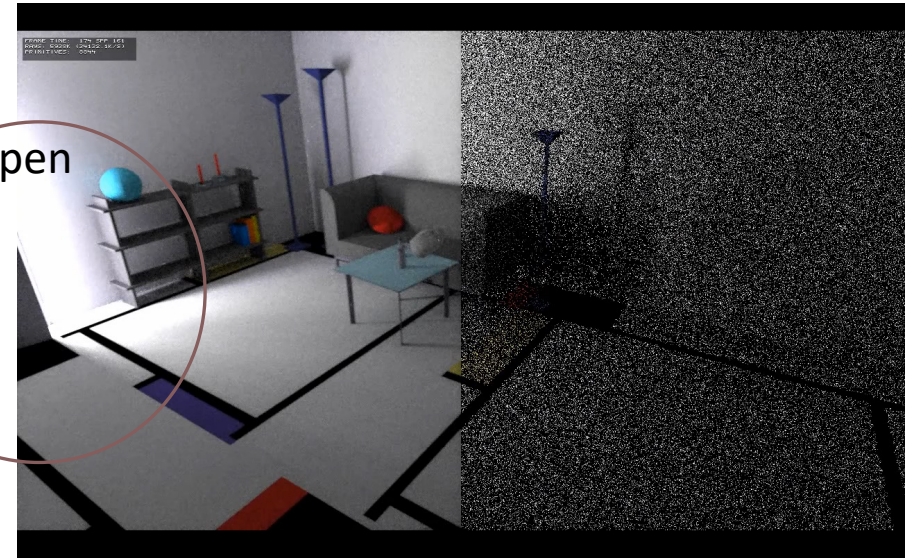
This idea inspired by a CGI rendering technique called Metropolis light transport.

- Computer animation often runs into a similar problem to us, where only a small fraction of light paths are detectable.
- Canonical example is a light source in another room that shines through a door that is only slightly cracked open.

CGI industry mainly renders scenes that are dominated by reflections.

- In IceCube, light transport is entirely scattering.

Slightly open door



Left: Rendering algorithm similar to Metropolis light transport.

Right: Standard path tracing algorithm.

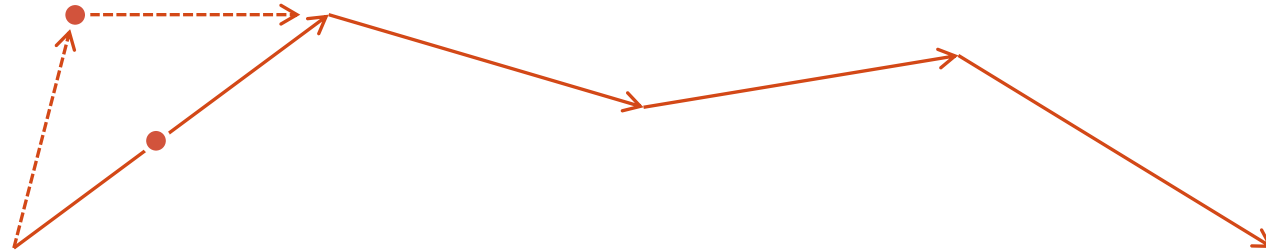


# Reversible jump MCMC

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The number of places where light scatters is not fixed.

- Thus, the dimensionality of the probability distribution is variable.






Reversible Jump Markov Chain Monte Carlo can change the number of dimensions in a probability distribution.

# Reversible jump MCMC

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The acceptance probability is based on the following ratio:

$$\frac{p_1(\vec{\phi})}{\boxed{p_0(\vec{\theta})q(\vec{q})}} \frac{P_{1 \rightarrow 0}}{P_{0 \rightarrow 1}} \left| \frac{\partial g(\vec{\theta}, \vec{q})}{\partial(\vec{\theta}, \vec{q})} \right| \quad \vec{\phi} = g(\vec{\theta}, \vec{q})$$

Padded probability distribution  Proposal rates  Jacobian of proposal function. 

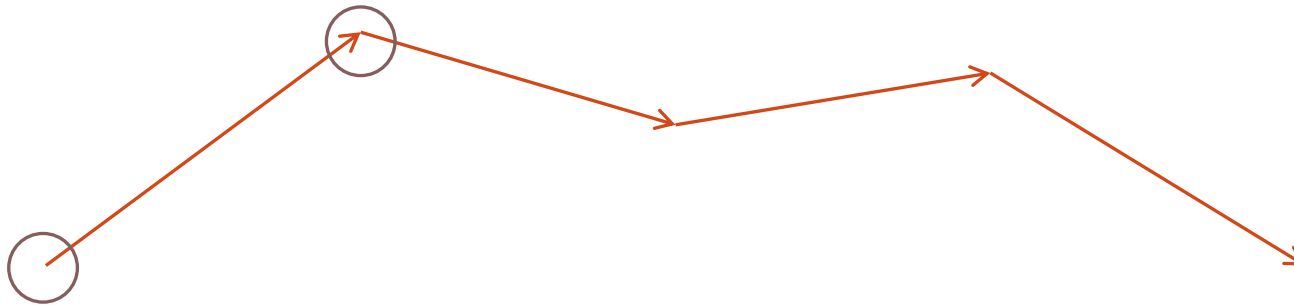
$q$  can be marginalised out later for free.

# Reversible jump for light propagation

A path with  $N$  vertices exists in  $\mathbb{R}^{3N}$

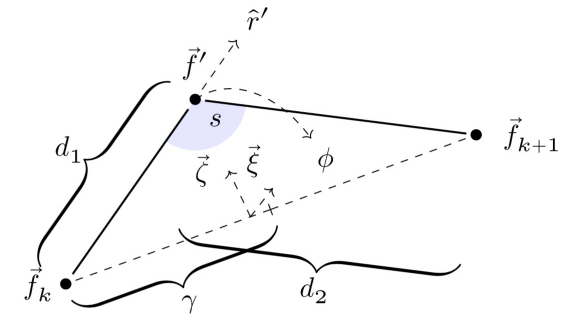
- We wish to propose a new path with  $N+1$  vertices.
- Requires a  $q$  with 3 parameters, and a choice of  $g$ .

$g$  selects a pair of vertices.



Then inserts a new vertex between them.

- Position of new vertex based on three random values from  $q$



# Path length distribution

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From the samples created by the MCMC, the probability distribution for path length can be easily extracted.

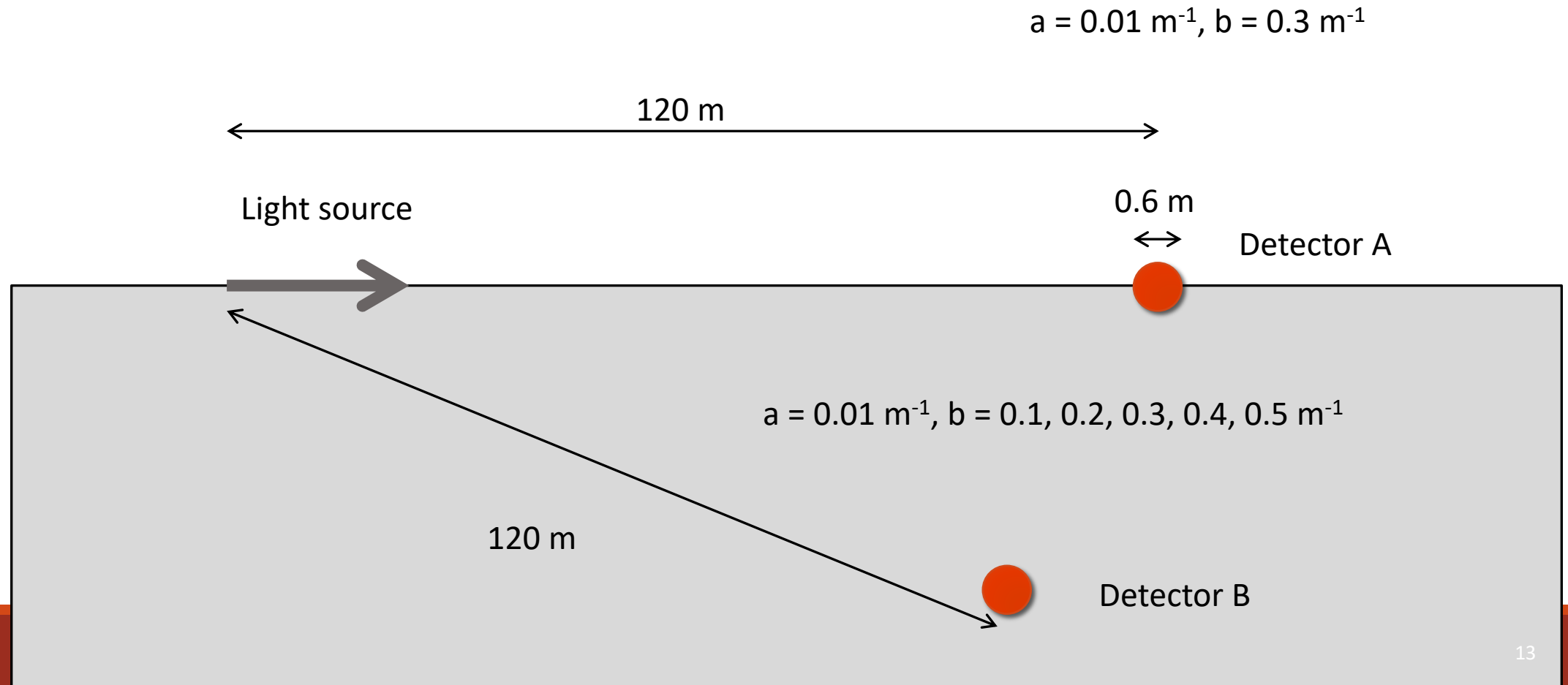
- IceCube measures photon arrival time, which is directly related to path length.
- $P(L < X)$  = fraction of samples where the length of the path is less than  $X$ .

To validate the method, the length distribution produced by the path sampler can be compared to one created using a ray tracer.

An MCMC usually requires a burn-in period, however this can be partially avoided by seeding the MCMC with the ray-tracer.

# Synthetic test case

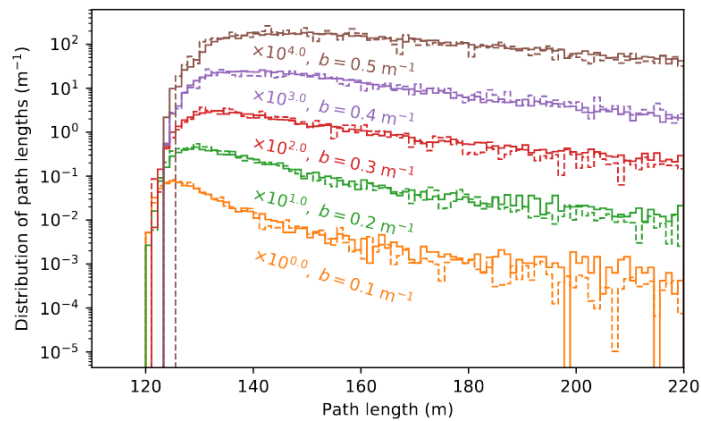
One light source, with two detectors



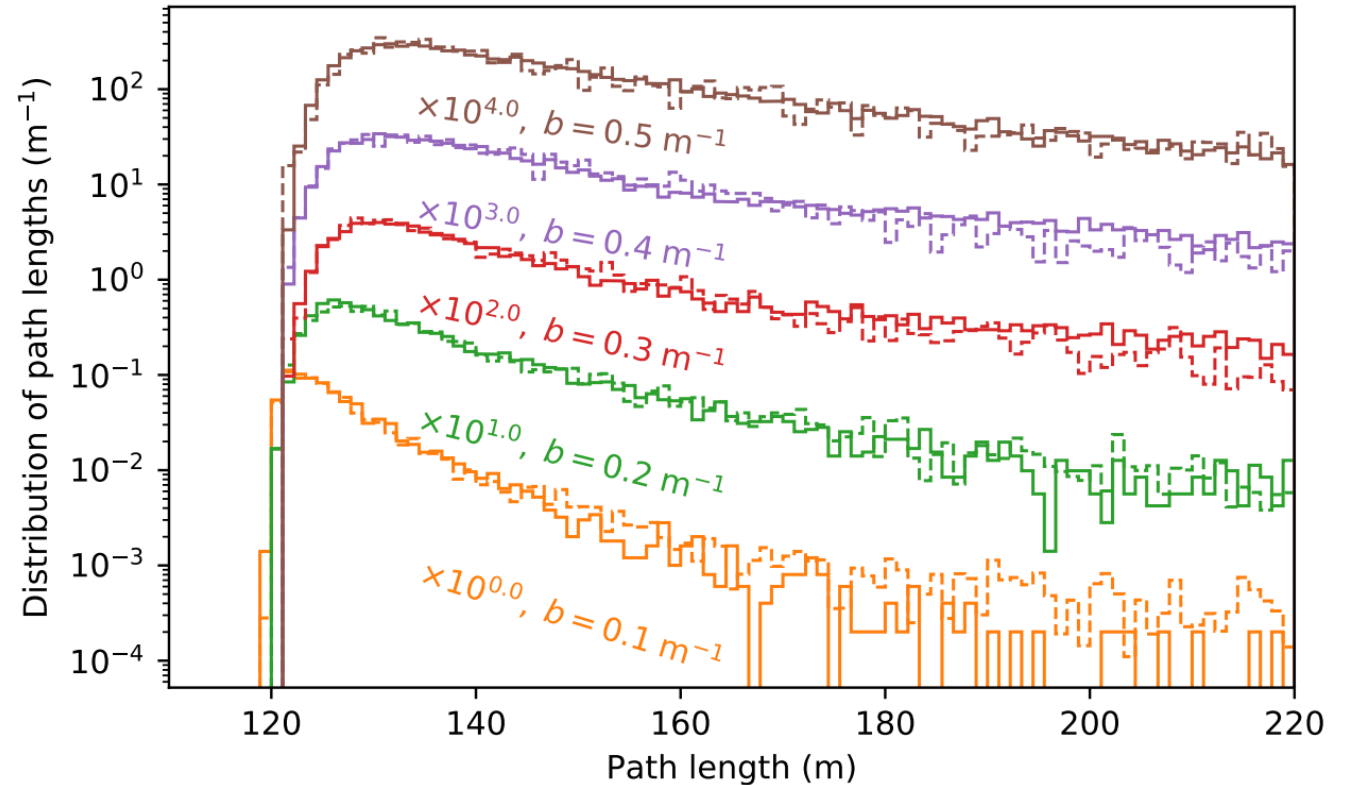
# Path length distribution

Solid: path sampler. Dashed:  
reference ray tracer

- Ray tracer was run until 5000 samples collected.
- Path sampler was run until results matched the path sampler.



Detector B



Detector A, Acceptance rate  $\sim 20\%$

Lines offset for readability

# Performance

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Ray tracer is also CPU based to allow a performance comparison.

<b>b</b>	<b>Ray tracer</b>	<b>Path sampler</b>
0.1 m <sup>-1</sup>	~46000 s	~23 s
0.2 m <sup>-1</sup>	~78000 s	~74 s
0.3 m <sup>-1</sup>	~99000 s	~232 s
0.4 m <sup>-1</sup>	~122000 s	~373 s
0.5 m <sup>-1</sup>	~156000 s	~416 s

Performance improvement of 300 to 1000 times faster.

- The  $b = 0.3$  to  $0.5 \text{ m}^{-1}$  cases are probably most comparable to conditions in IceCube.

# Other applications

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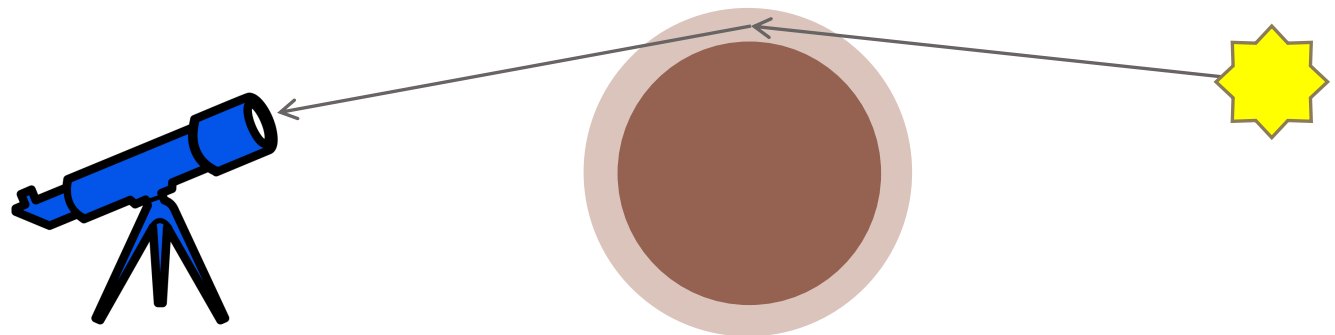
This approach to simulation is useful when initial and final states are highly constrained.

Litmus test:

- Are you throwing out the vast majority of your events (99.9%+) due to them not meeting one of these constraints?

Constraints do not have to just be in position.

- Eg: initial and final angle for light passing through a planetary atmosphere.





# Other applications

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Path does not just have to describe light.

- Eg: Simulation of transport of neutrons.

Constraints could be discrete parameters.

- Eg: Simulation of atmospheric showers.
  - Initial condition: particle must be a nucleus.
  - Final condition: shower products must reach underground detector.

May also be possible to incorporate selection cuts into the constraint.

# Disadvantages

Relative and especially absolute light yields are difficult to calculate.

Relative light yield is given by the ratio of normalisations for each detector.

- This is equivalent to finding a Bayes factor in Bayesian inference.

We can use the geometric estimator:

$$\mathcal{B} = \frac{\mathbb{E}_A \left[ \sqrt{p_B(x) / p_A(x)} \right]}{\mathbb{E}_B \left[ \sqrt{p_A(x) / p_B(x)} \right]}$$

To estimate the variance on the light yield

- Ray tracer and path sampler were run four times. (standard deviation in parentheses.)

<b>b</b>	<b>Ray tracer</b>	<b>Path sampler</b>
0.1 m <sup>-1</sup>	0.67(3)	0.64(7)
0.2 m <sup>-1</sup>	0.79(2)	0.77(1)
0.3 m <sup>-1</sup>	0.73(2)	0.76(7)
0.4 m <sup>-1</sup>	0.68(2)	0.64(4)
0.5 m <sup>-1</sup>	0.59(2)	0.50(6)

- Path sampler has higher variance.

# A new approach?

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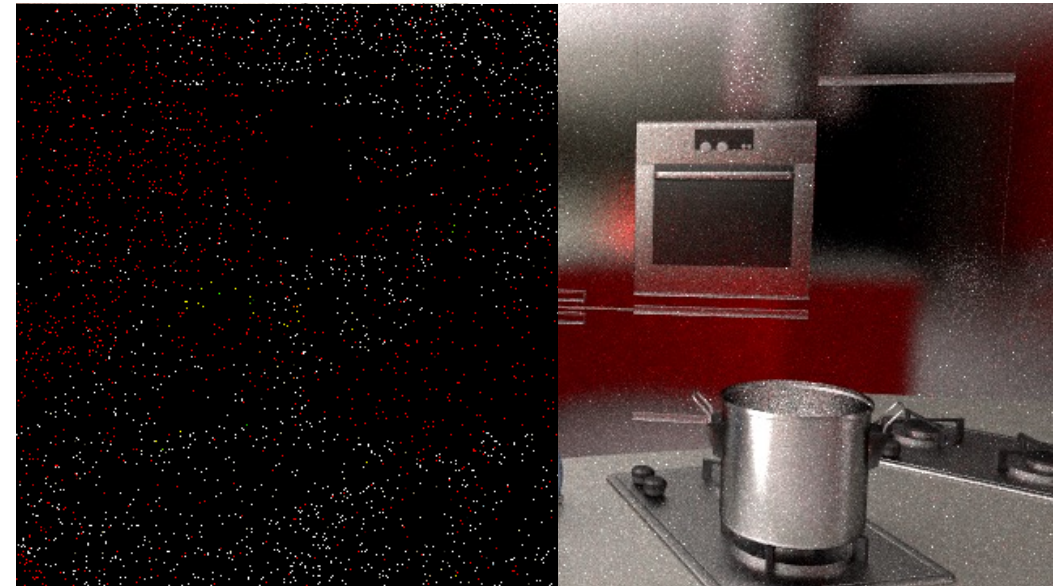
Can we combine the advantages of path tracing and metropolis light transport?

- Perhaps with Guided Path Tracing.

Run path tracing in batches.

- Use the “useful” rays in each batch to bias the next batch.
- Results in minimal biasing, and therefore minimal variance.
- Bias can be corrected for like event-weighting.

A further advantage is that this can slot into many existing physics simulations.



Left: Path tracing. Right: Guided path tracing.

# Conclusion

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Simulation of light can be posed as a path integral from which samples can be drawn.

Reproduces the timing distribution of light incident on a detector.

- Up to 1000 times faster than a ray tracer in synthetic test case.

Method is generally applicable to a wide range of problems.

- When initial and final states are highly constrained.

# Backup

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# Angular scattering distribution

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Distribution

$$\sigma(\cos \psi) = f_{\text{SL}} p_{\text{SL}}(\cos \psi) + (1 - f_{\text{SL}}) p_{\text{HG}}(\cos \psi),$$

Simplified Liu:

$$p_{\text{SL}}(\cos \theta) = \frac{1}{2} \frac{1+g}{1-g} \left[ \frac{1+\cos \theta}{2} \right]^{\frac{2g}{1-g}}$$

Henyey-Greenstein:

$$p_{\text{HG}}(\cos \theta) = \frac{1}{2} \frac{1-g^2}{(1+g^2-2g\cos \theta)^{3/2}}$$

# Detection probability

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Conditional detection probability:

$$\rho(\vec{f}_{n-1}) = \exp(3 \cos \omega - \ln \cosh(2 \cos \omega + 0.7) - 1)$$

$$\cos \omega = \hat{\rho} \cdot \frac{\vec{f}_{n-1} - \vec{\eta}}{|\vec{f}_{n-1} - \vec{\eta}|}$$

Chosen to follow IceCube DOM angular response.

# Jump distributions

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$$q(s) = \frac{\beta e^{-\beta \cos s}}{2 \sinh \beta} \sin s,$$
$$q(t) = (2 + 2 \cosh t)^{-1},$$
$$q(\phi) = \frac{1}{2\pi},$$



# Jump rates

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$$p(n \rightarrow n + 1, k) = \frac{\tau_b(k)}{\sum_{l=1}^{n-1} \tau_b(l)},$$

$$p(n \rightarrow n - 1, k) = \frac{1}{n - 2}.$$

# Incident angle distribution

For a smoothly varying  $b$ :

