



cherenkov
telescope
array

The Cherenkov Telescope Array: an observatory perspective

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CTA: the Cherenkov Telescope Array

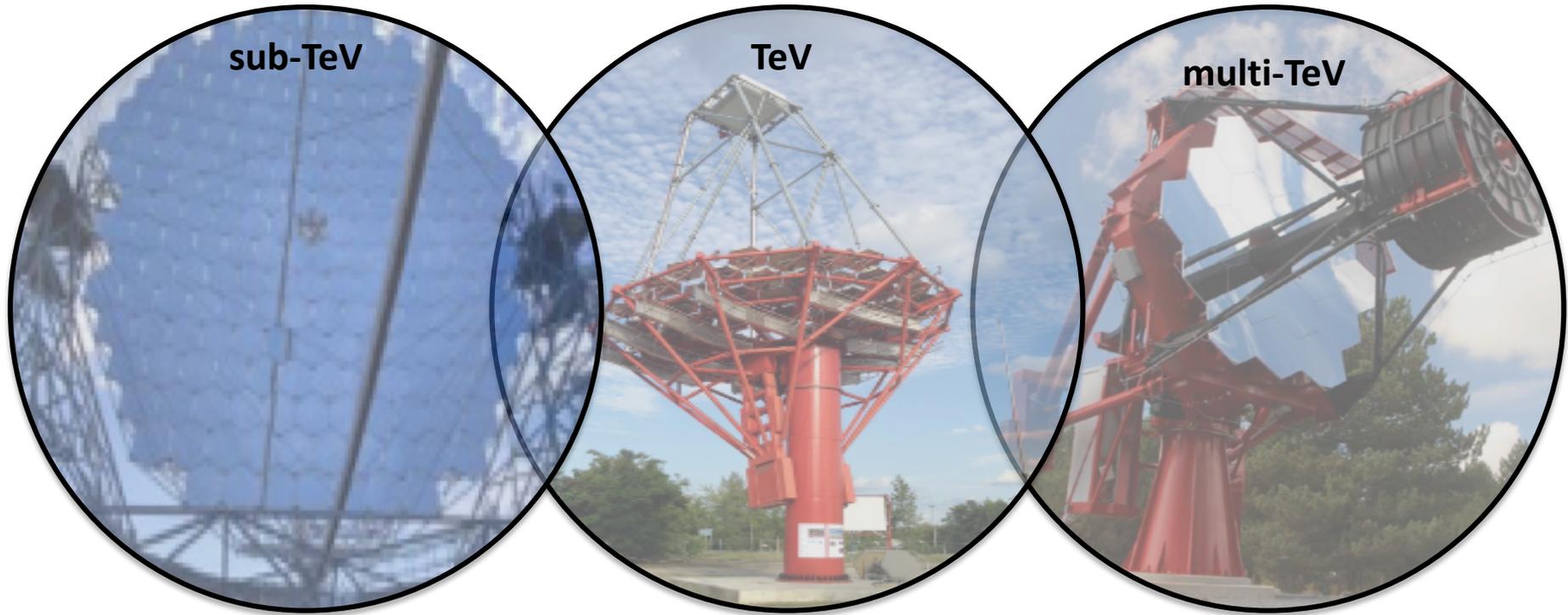
Next generation ground-based gamma-ray observatory



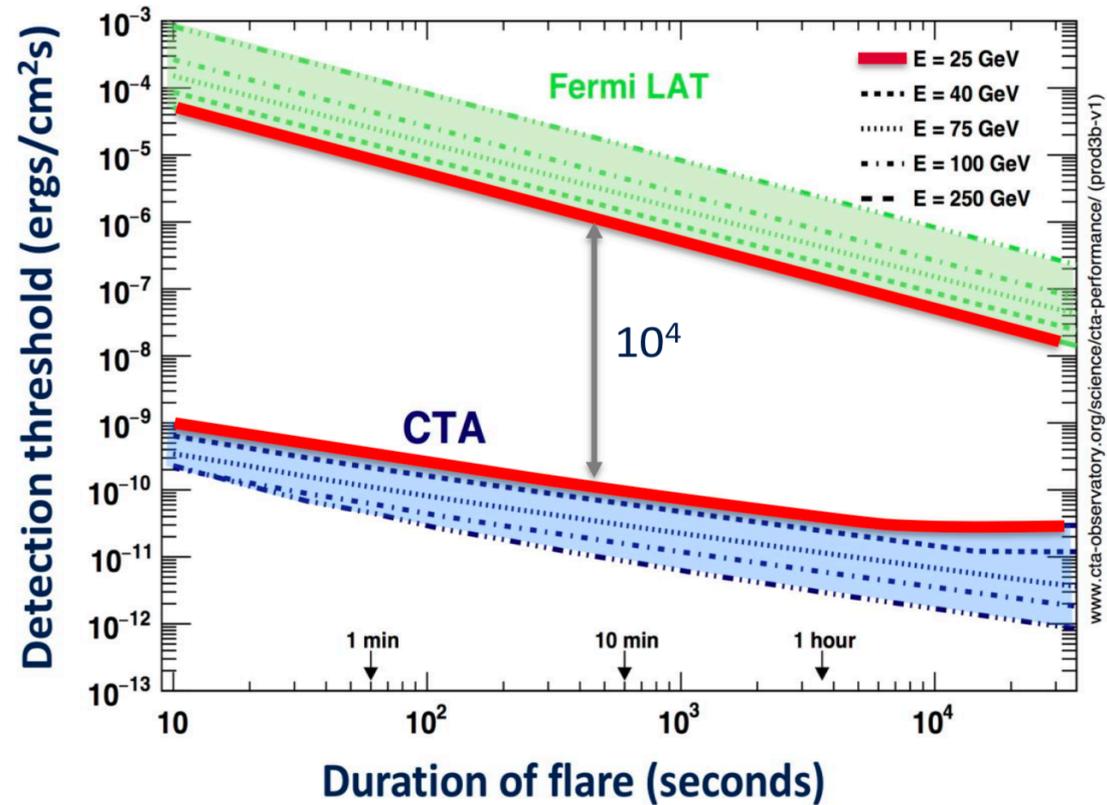
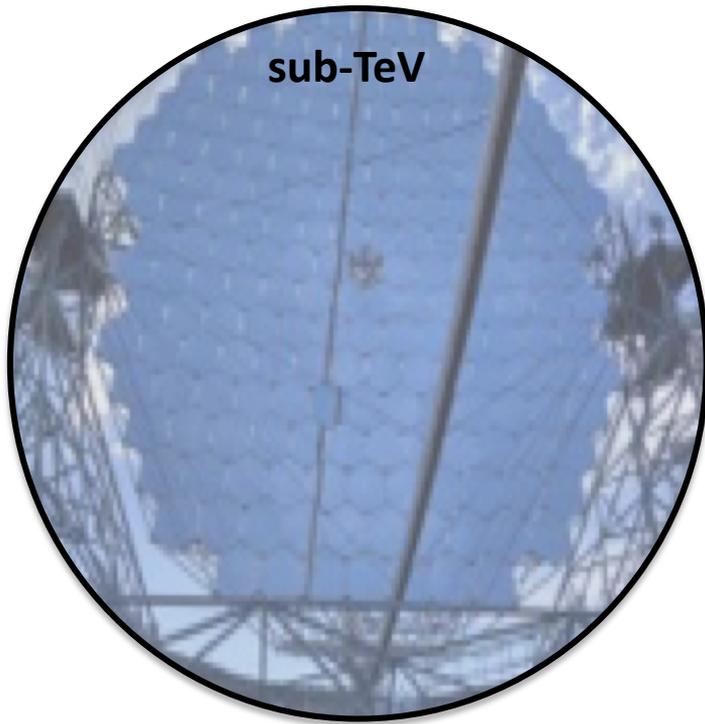
- Science cases and design drivers
- Performances
- Observatory & operations
 - Operations in a MWL/MM era



Science cases and design

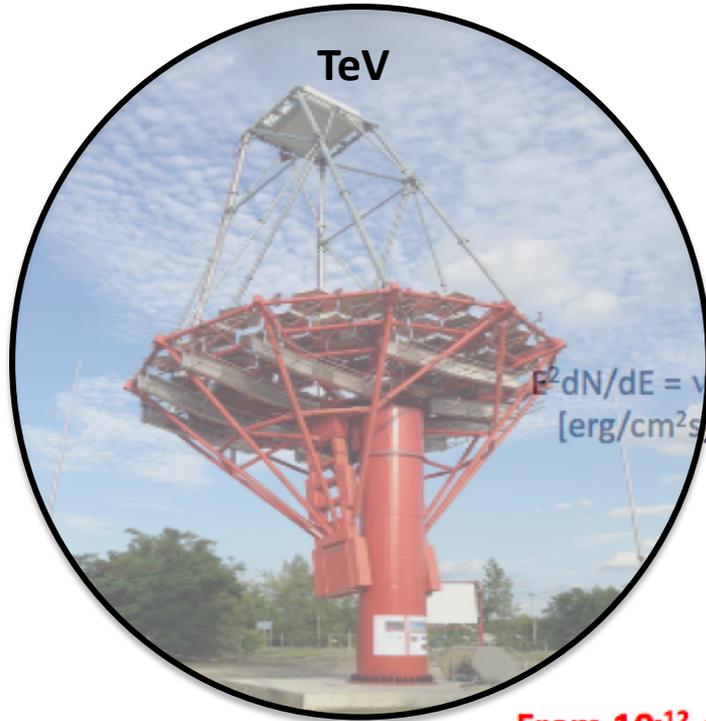


sub-TeV energies - Flare sensitivity

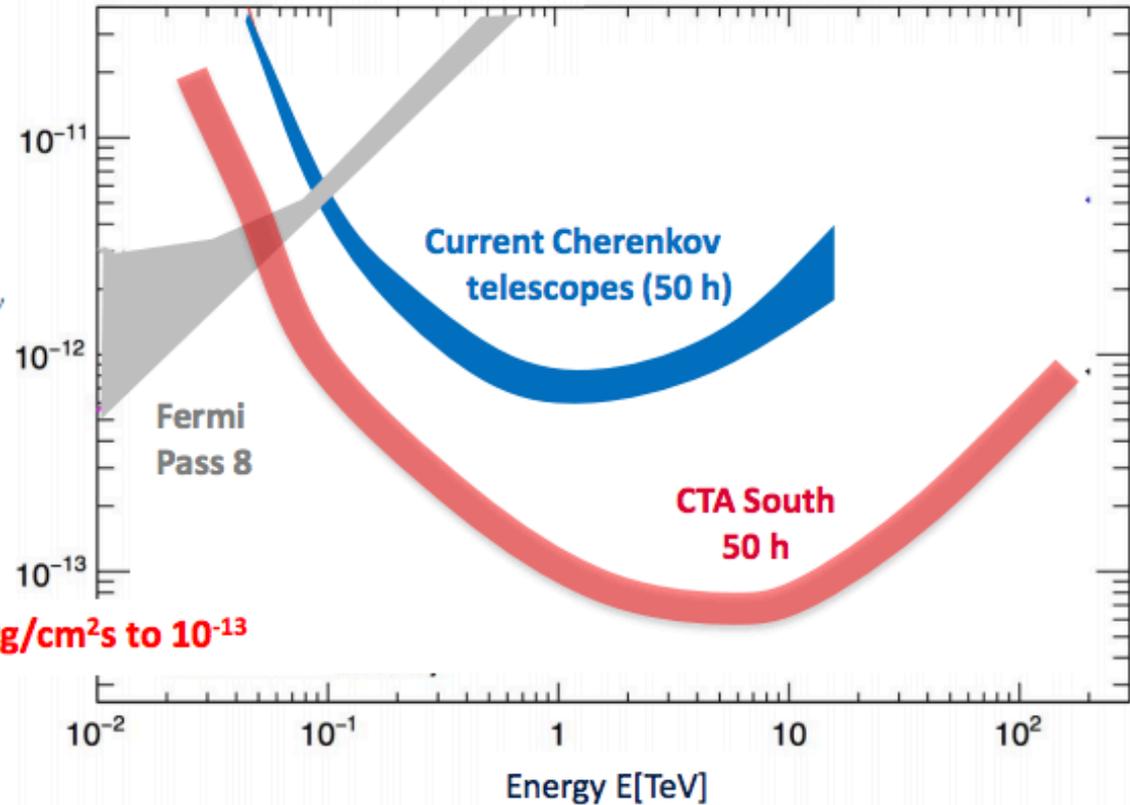


- Deepest sensitivity for short timescale phenomena
- Time domain unexplored
→ cosmological sources

TeV energies - Sensitivity (steady sources)



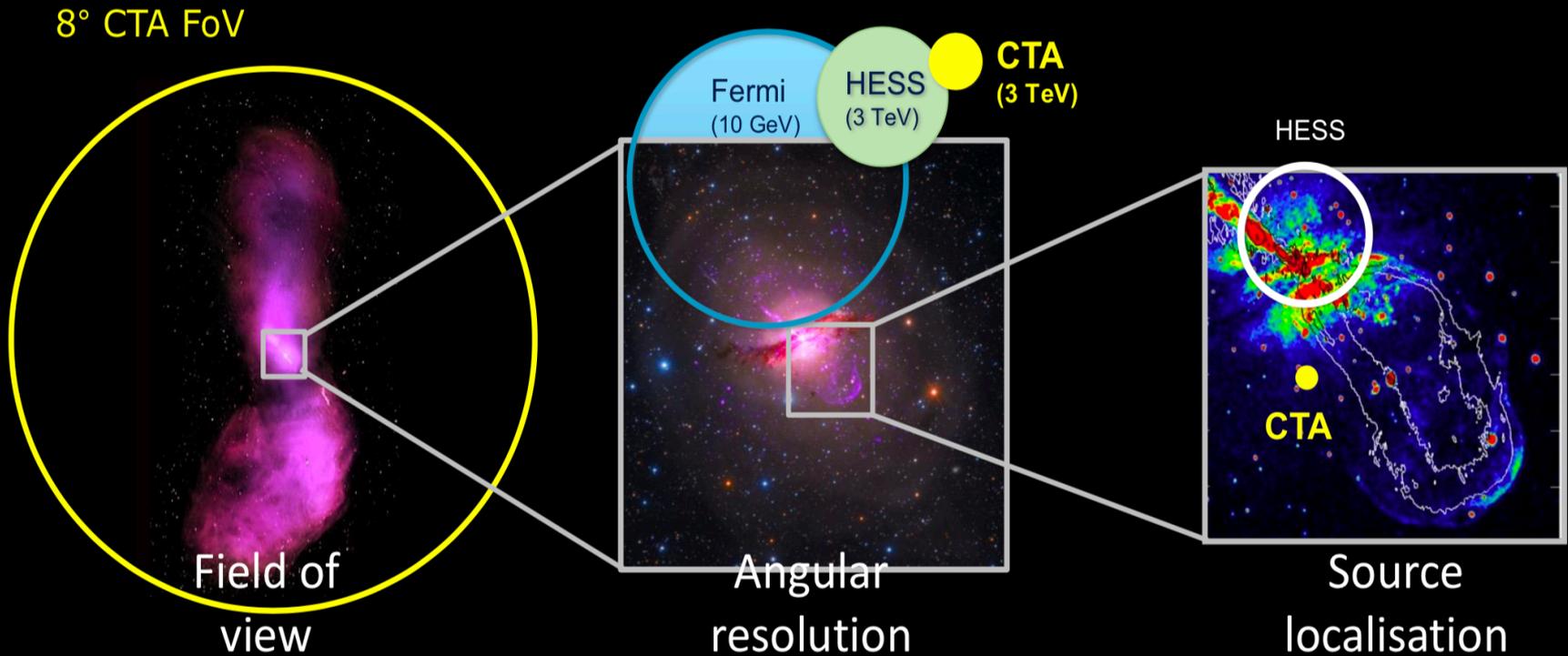
From 10^{-12} erg/cm²s to 10^{-13}



- deepest sensitivity ever
- arcmin angular resolution
- large FoV

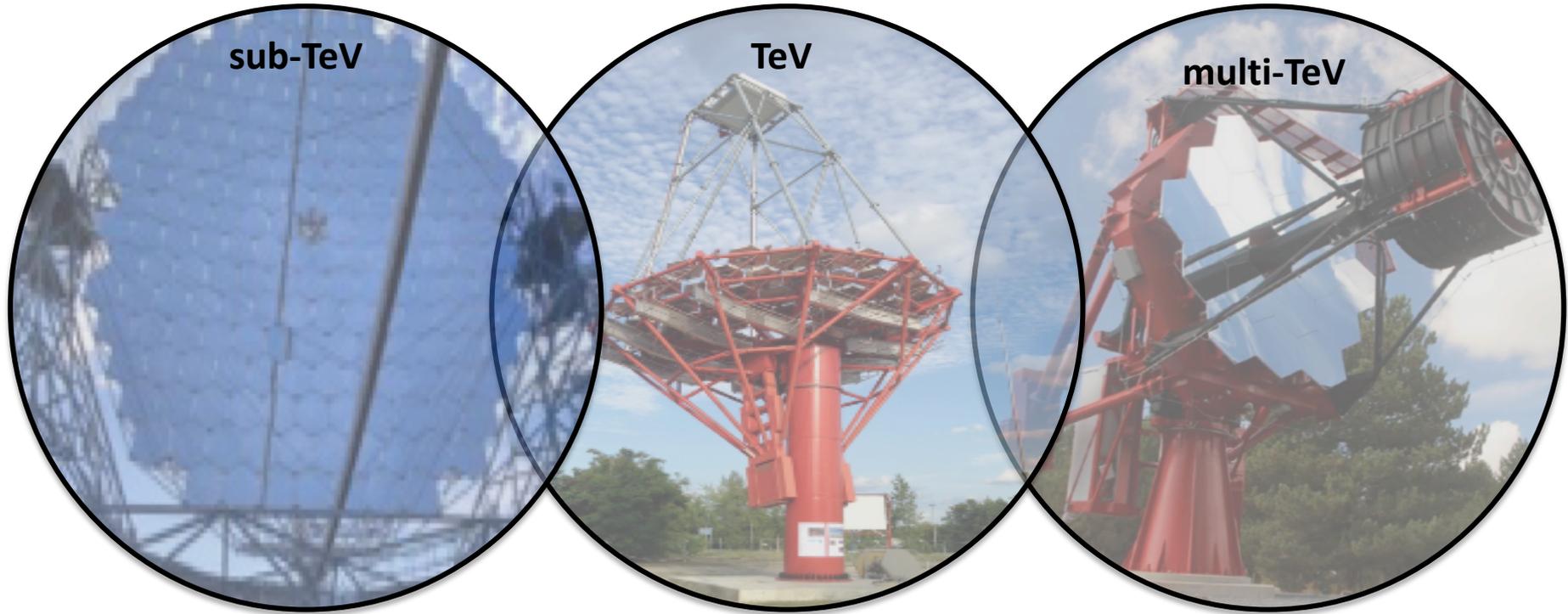
- Surveys & precision studies

TeV/multi-TeV energies - Resolving power



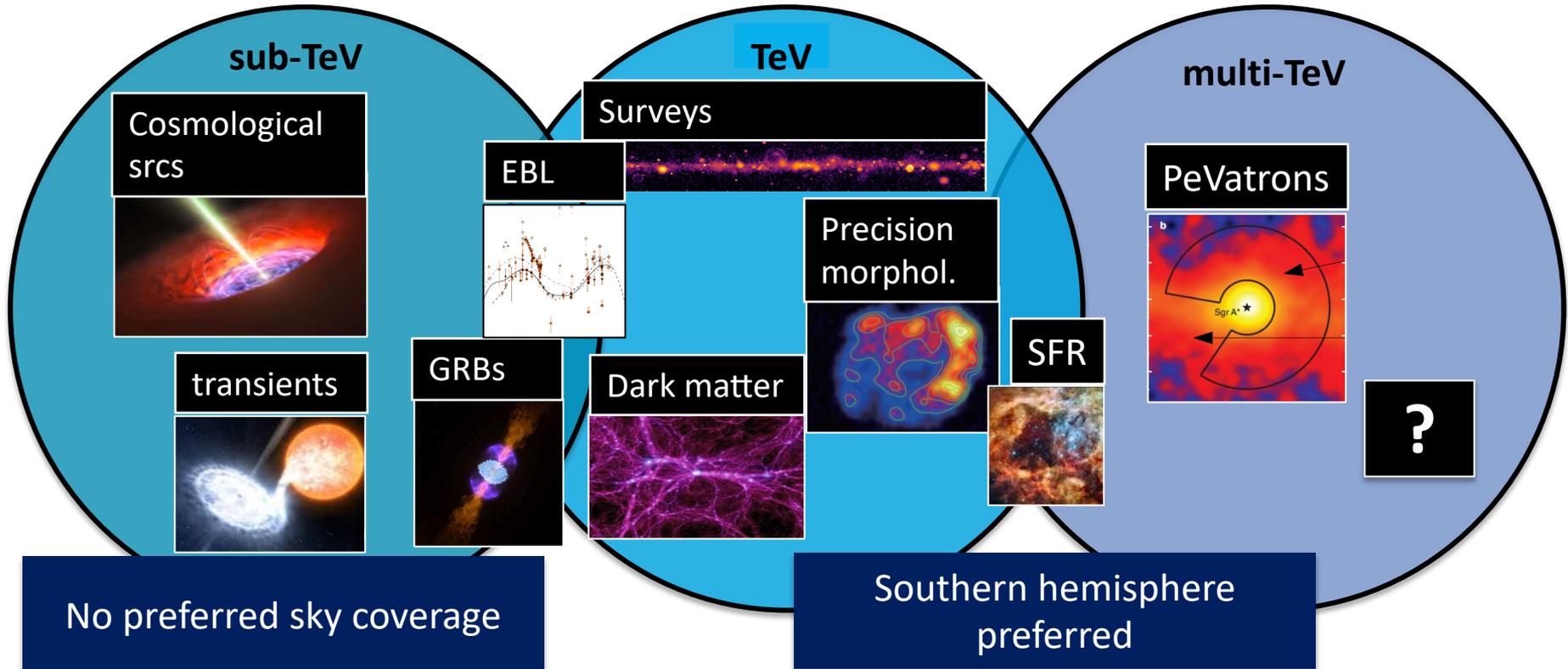
Example: nearby active galaxy Centaurus A

Science cases and design



- Deepest sensitivity for short timescale phenomena
- Time domain unexplored → cosmological sources
- deepest sensitivity ever
- arcmin angular resolution
- large FoV
- Surveys & precision studies
- Precision measurements in a still little explored energy range
- 100 TeV range unexplored
- precision studies

Science cases



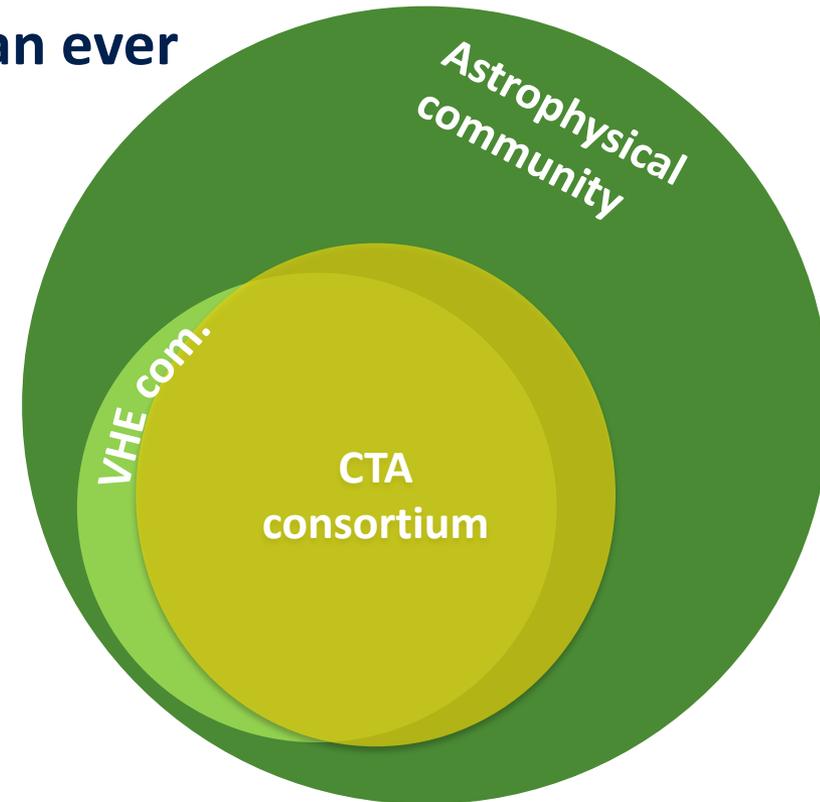
- **Mainly CTA consortium involved in the definition of the science cases**

(Science with CTA, CTA Consortium 2019 - <https://doi.org/10.1142/10986>)

- **The science case is stronger than ever**
 - γ rays – neutrinos correlation as proof for the mechanisms at work in the most violent explosions in our Universe
 - Detection/non-detection of the electromagnetic counterparts of the GW
 - Pulsars as a potential new population of TeV emitters

- **The science case is stronger than ever**

- γ rays – neutrinos correlation as proof for the mechanisms at work in the most violent explosions in our Universe
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- **The advent of the multi-messenger has enlarged the interest in the overall astronomical community on the VHE astrophysics**

- **The first ground-based gamma-ray observatory**
 - serve large user community in an open and fair way
- **30 yr of lifetime**
 - Significant effort for maintenance and operations costs optimization
- **One legal entity: CTAO GmbH in the process to become an ERIC with HQ in Bologna (Italy)**
- **Two Telescope arrays, one Observatory**
 - Inter-site coordination
 - Uniform approach to scientific ops
- **The Science Data Mgmt Center in Zeuthen (Germany)**
 - CTA is a software instrument

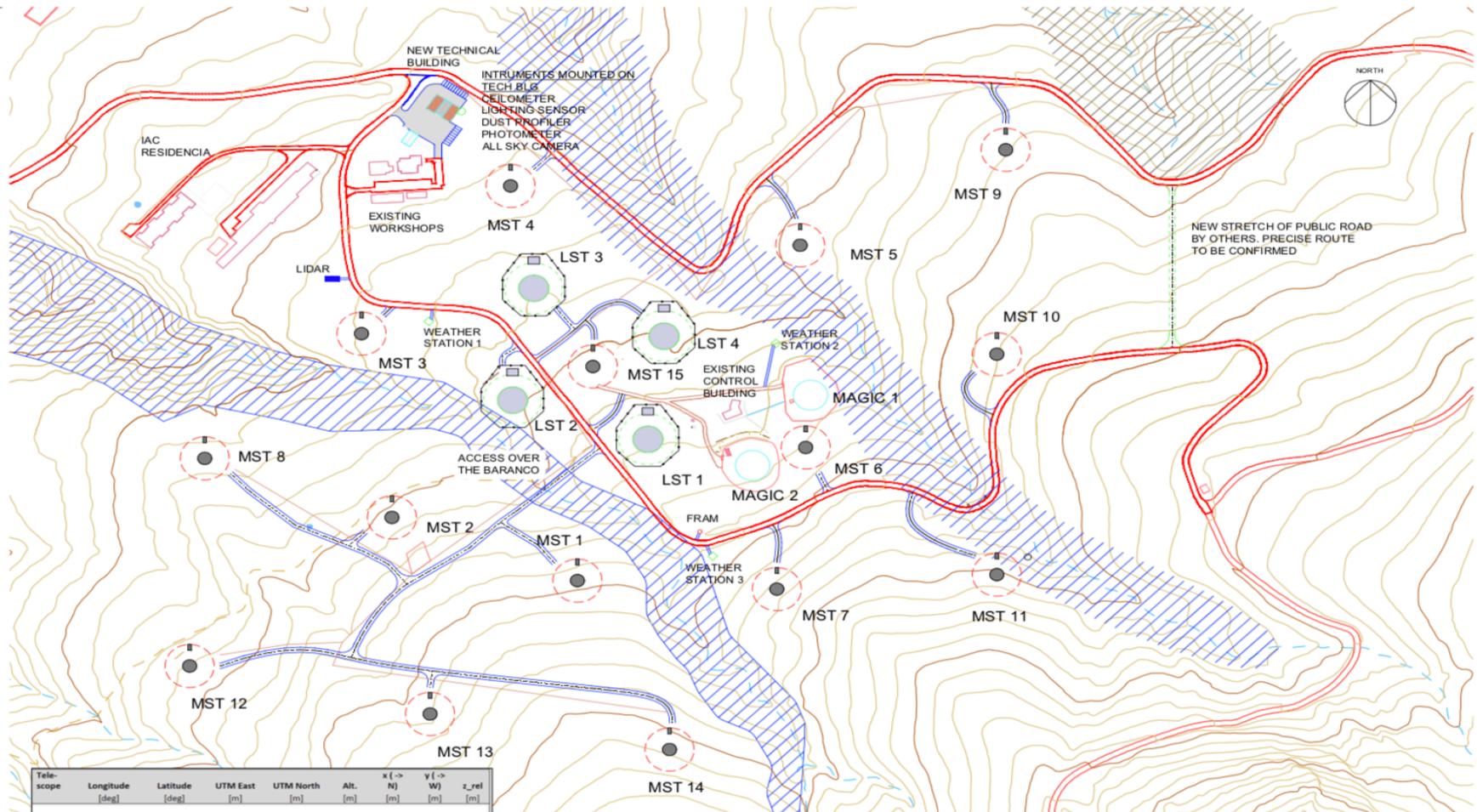
CTA Observatory



CTA-North site



- 4 LSTs + 15 MSTs (baseline configuration)
 - Focus on sub-TeV and TeV energy range



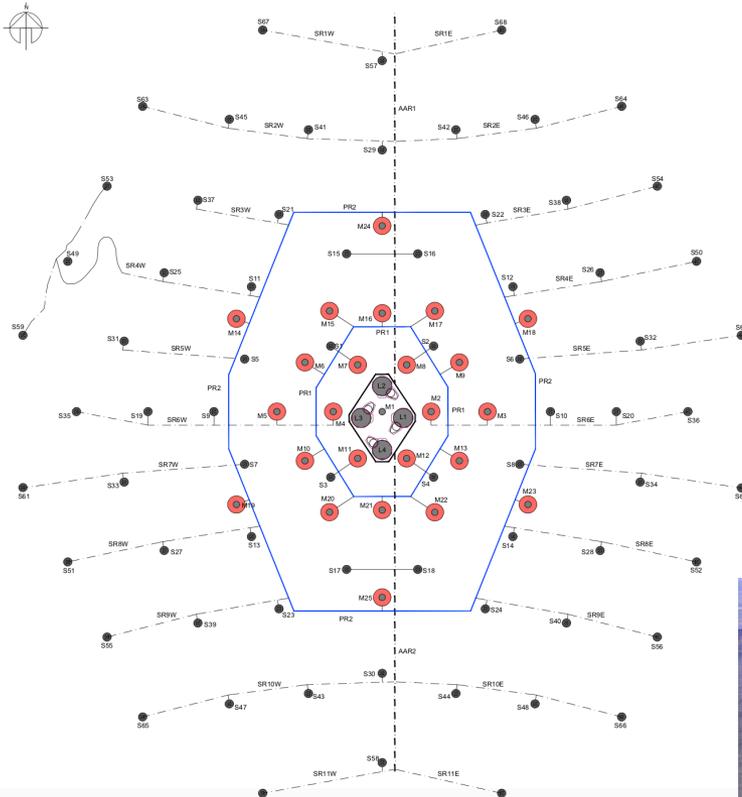
CTA-North site



LST1



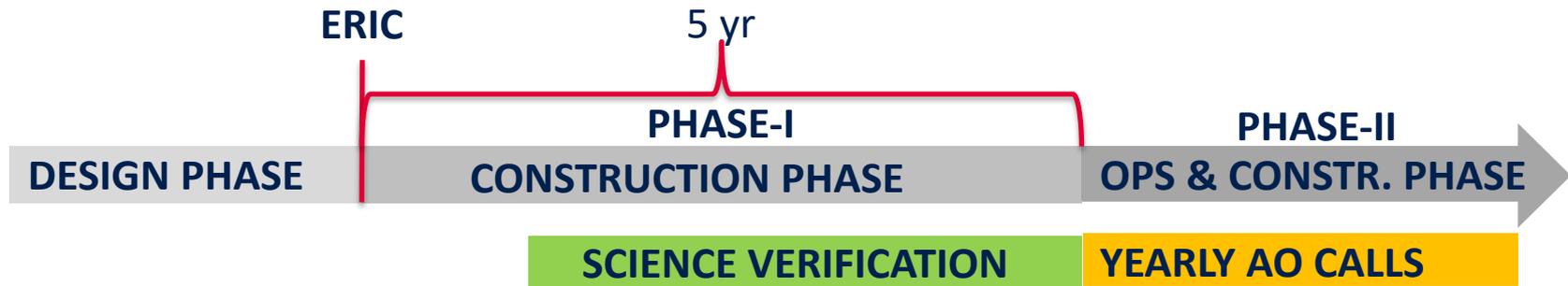
- 4 LSTs + 25 MSTs + 70 SSTs (baseline-configuration)



- Site agreement signed in Dec 2018
- Aim to start with site infrastructure construction soon



Project Status



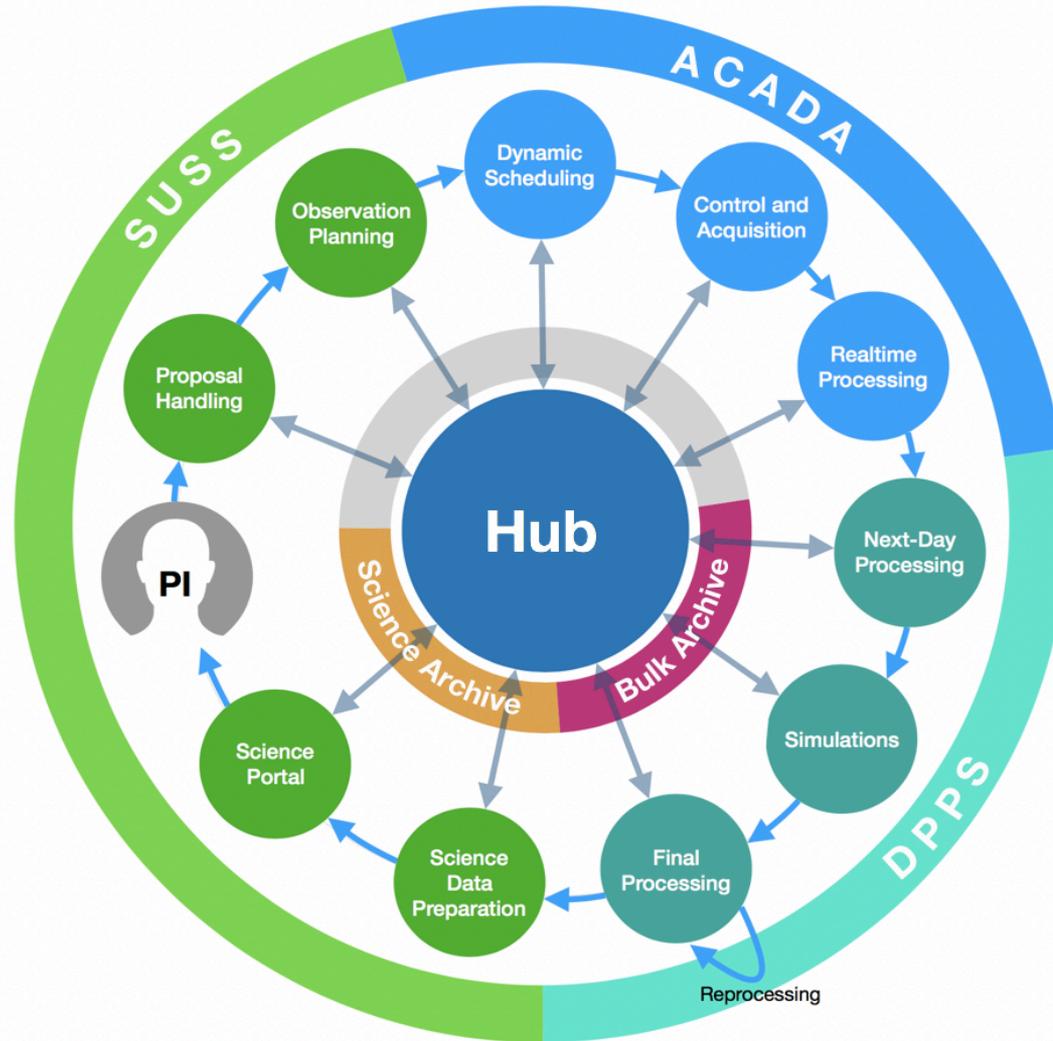
- **Phase-I: construction of a reduced-baseline configuration**
 - a significant performance improvement wrt the currently running facilities
 - guarantees high-impact science covering most of the science cases
 - guarantees a significant increase of the discovery space
- **Phase-II: operation of the phase-I configuration + construction towards the final baseline**

A proposal-driven observatory

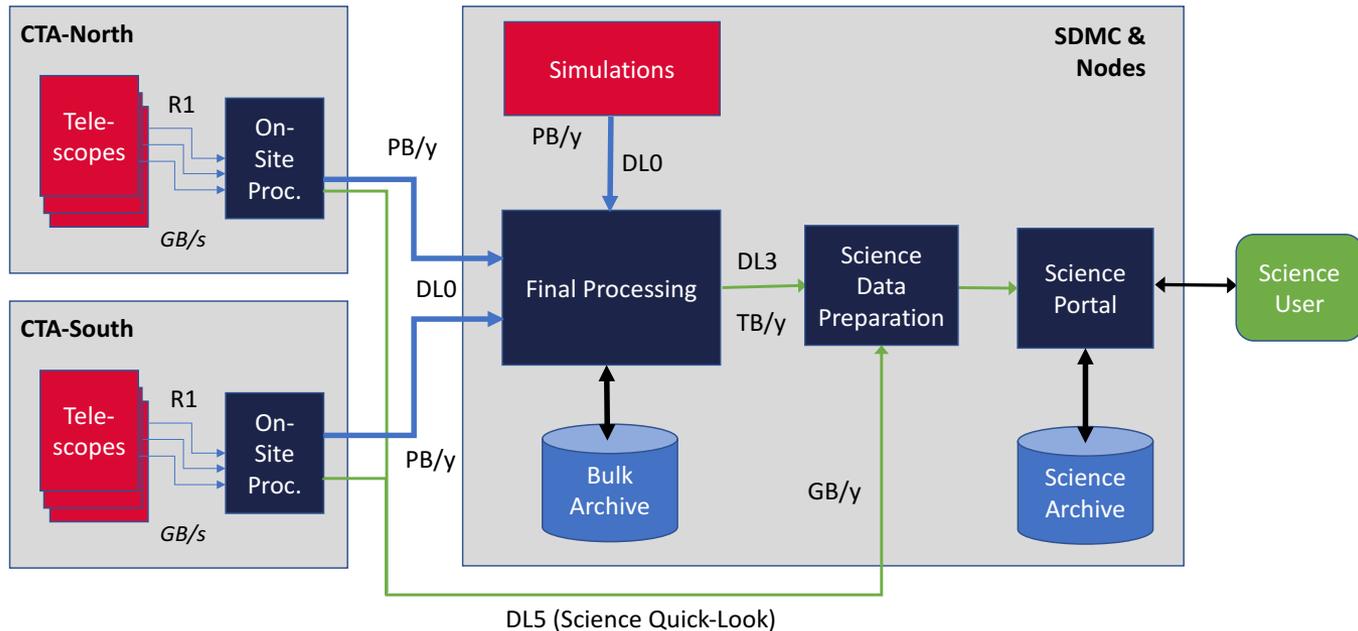


- **Announcement of opportunity call every year**
 - **Proposal evaluation through a peer-review process**
 - **Each proposal is associated to one unique PI**
 - **After 1yr proprietary period data are open**
 - **Different data quality levels corresponding to distinct observing conditions**
- **Scientific proposals are classified in:**
 - **Regular PI proposals**
 - Tens to hundreds of hours of observation
 - **Key Science Projects:** large observation programs (from hundreds to thousands hr) that need to be addressed in a coherent fashion generating legacy data sets for the community
 - 40% of observation time devoted to KSPs for the first 10 yr
 - **Discretionary Director's Time (DDTs) proposals**

Overview on Science Operation Processes



Data Flow & Analysis Categories



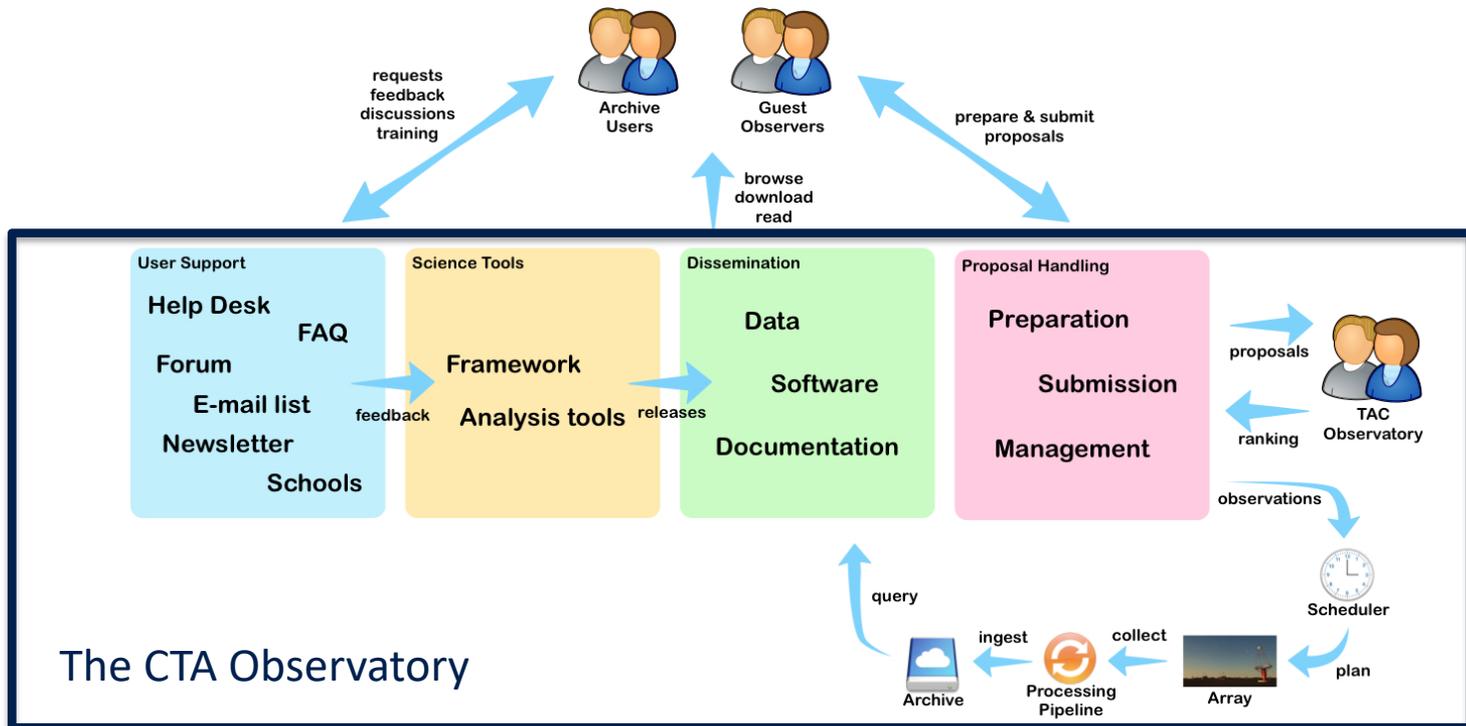
Three analysis categories

- **Category A:** at the sites, **real-time**, generate internal science alerts
- **Category B:** at the sites, offline (e.g. next day) → Higher sensitivity than category A
- **Category C:** offsite, final analysis results with maximum sensitivity

Science User Perspective



- **Science User get access to the analysis results**
 - on different timescales depending on the type of analysis
- **Products and services are in the core of the CTA Science Platform**
 - Access to results for PIs only via the CTA Science Platform



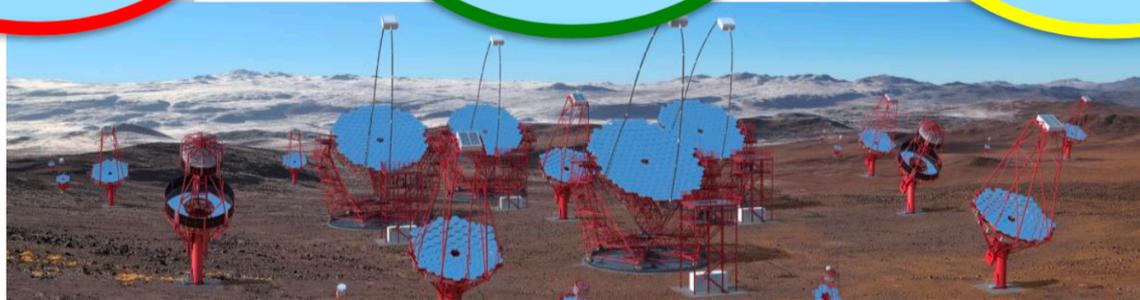
Crucial: MWL & MM links



Target selection & ToOs

Object properties

Wide-band / MM SED



- **Catalogs for source associations**
 - Radio, IR, optical, X-rays, gamma rays
 - State-of-art catalog of the currently running VHE facilities
- **Radio surveys for the study of the diffuse emission**
- **Specific (non-flaring) source archival data**
 - Crucial to disseminate our data in a standard format to favor the interoperability and MWL/MM analyses

MWL/MM observational campaigns



- **MWL/MM campaigns are organized for**
 - Typical timescales days to months
 - expected exceptional events (periodic sources, transitioning sources, long-duration flaring states)
 - correlation studies (i.e. giant pulses, FRB counterparts ..)
- **They require fixed predefined timeslots**
 - policy defined by previous agreements
- Constant communication between the involved facilities to react to possible failures of the partners (bad weathers, technical problems...)

Target of Opportunity Observations



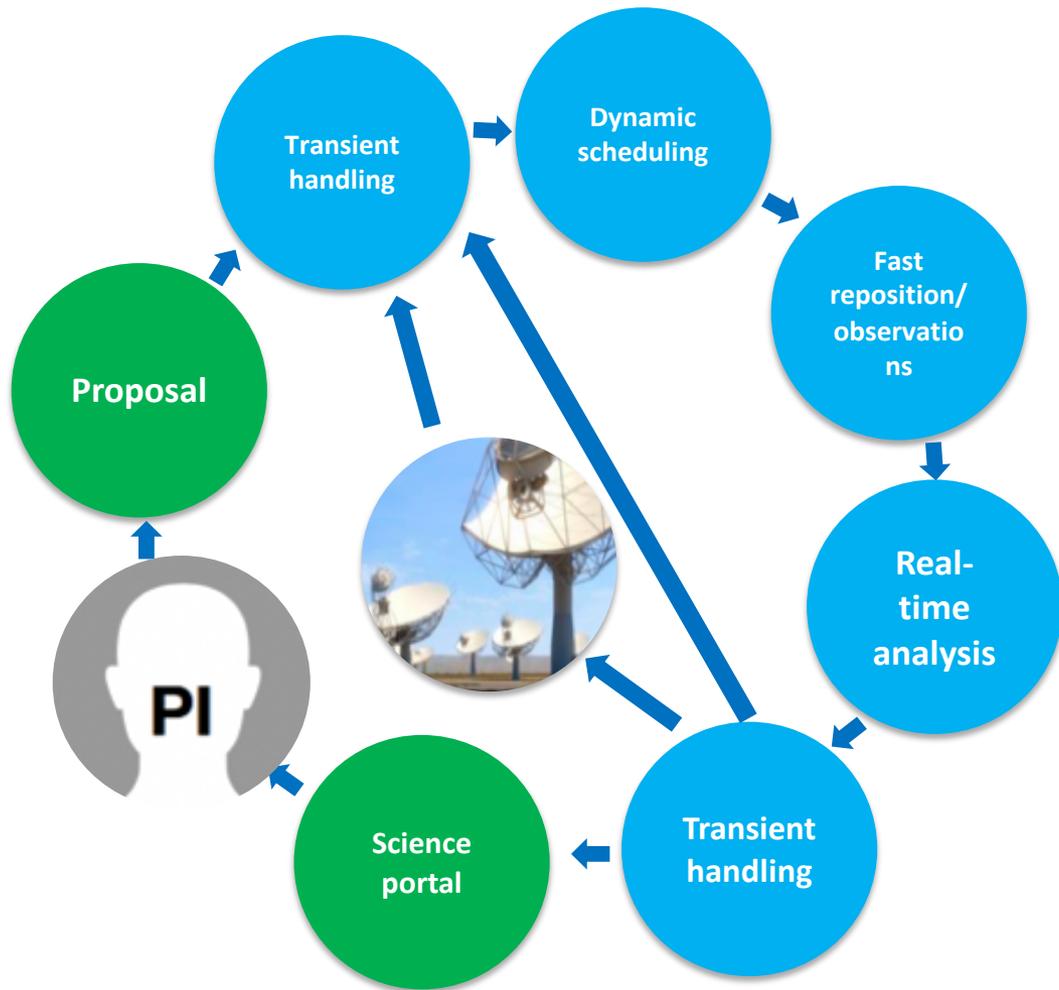
- **Monitoring ToOs**
 - Typical scheduling timescales **days to weeks**
 - **Monitoring** can be performed
 - **by different facilities** (it requires previous agreements)
 - **by a sub-array of CTA telescopes**
- **Offline ToOs**
 - Typical scheduling timescales **days to weeks**
 - **Request** to activate the corresponding proposal **sent on day timescale by the PI to the CTAO**, that schedules the observations off-site/off-line

Target of Opportunity Observations



- **Externally generated real-time ToOs (GRBs, SGRs, Novae, AGNs...)**
 - **Automatic handling of alerts on ~second timescales**
 - **The alert system shall have a memory of the alerts during day-time**
 - **Alert prioritization will be a must in the era of transient factories**
- **Internally generated real-time ToOs**
 - **Serendipitous detection in the FoV detected by the real-time analysis on ~minute timescales**
 - **Fast communication of the results to the other Observatories to enable follow-up observations**
 - **expected rate low, but identified events with high impact**

CTA as a transient facility



- **handling of any reaction to internal and external alerts by CTAO**
 - All functions planned to be compatible with **international standards to ensure interoperability**

Synergies with other facilities



- **MoU CTAO-SKAO about to be signed**
 - Exchange of internal documents (data policies, architectures, ops concepts,...)
 - White paper CTA-SKA synergies
- **MoU CTAO-EGO signed in 2016**
- A series of white/yellow papers exploring the synergies with future observatories
 - Theseus
 - Athena

