



# Muon Calibration in CTApipe

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# CTA Muon Calibration

- Telescope / Camera teams must:
  - Be able to trigger on incoming muons
  - Flag candidate muon events
  - i.e. perform muon preselection
- Telescope / Camera teams may:
  - Develop algorithms for internal use
  - Perform own simulations e.g. with degraded efficiency
- BUT: all telescopes will be calibrated by the same analysis and calibration chain.
- Muon software of teams beyond pre-selection will NOT be used in standard muon calibration → **ctapipe**



# Muon Calibration Algorithms

- Many algorithms exist and are currently in use
- Ideally, one algorithm for all telescopes / cameras
- Which one should be used?
  - The one that performs the best!
- How to know? Compare algorithms on figures of merit in same software framework → ctapipe
- Recommend: If you think an algorithm could be \*the one\*, it should be implemented in ctapipe



# Muon Calibration (reconstruction)

- Flagged candidate muon events enter. Two step procedure:
- Circle fitting
  - Chaudhuri-Kundu analytical solution → *implemented*
  - Chernov-Osokov analytical solution
  - Hough Transform
  - Karimäki chi-square fit
  - Taubin fit
  - Kasa method
- Intensity fitting
  - Chi-square fit to 1D profile
  - Chi-square fit to 2D image
  - Pixel-wise Log-likelihood fit to 2D image → *implemented*
- Per event fit parameters stored



# Storage Classes: MuonRingParameter

- Container classes in io/containers.py
- After ring/circle fitting, store parameters:
  - run\_id, event\_id, tel\_id, inputfile
  - ring\_center\_x, ring\_center\_y,
  - ring\_radius
  - ring\_phi, ring\_inclination
  - ring\_chi2\_fit, ring\_cov\_matrix
  - ring\_fit\_method



# Storage Classes: MuonRingParameter

- Container classes in `io/containers.py`
- After ring/circle fitting, store parameters:
  - `run_id`, `event_id`, `tel_id`, `inputfile`
  - `ring_center_x`, `ring_center_y`,
  - `ring_radius`
  - `ring_phi`, `ring_inclination`
  - `ring_chi2_fit`, `ring_cov_matrix`
  - `ring_fit_method`
- Example ring fitting:
  - `image/muon/muon_ring_finder.py`
  - `ChaudhuriKunduRingFitter` class



# Storage Classes: MuonIntensityParameter

- Container classes in io/containers.py
- After intensity fitting, store parameters:
  - run\_id, event\_id, tel\_id, inputfile
  - ring\_completeness, ring\_num\_pixel, ring\_size, off\_ring\_size,
  - ring\_width, ring\_time\_width,
  - intensity\_cov\_matrix, COG\_x, COG\_y
  - impact\_parameter, impact\_parameter\_chi2
  - impact\_parameter\_pos\_x, impact\_parameter\_pos\_y
  - optical\_efficiency\_muon, mask, prediction,
  - intensity\_fit\_method



# Storage Classes: MuonIntensityParameter

- Container classes in io/containers.py
- After intensity fitting, store parameters:
  - `run_id`, `event_id`, `tel_id`, `inputfile`
  - `ring_completeness`, `ring_num_pixel`, `ring_size`, `off_ring_size`,
  - `ring_width`, `ring_time_width`,
  - `intensity_cov_matrix`, `COG_x`, `COG_y`
  - `impact_parameter`, `impact_parameter_chi2`
  - `impact_parameter_pos_x`, `impact_parameter_pos_y`
  - `optical_efficiency_muon`, `mask`, `prediction`,
  - `intensity_fit_method`
- Example intensity fitting:
  - `image/muon/muon_integrator.py` : `MuonLineIntegrate` class



# Progress so far

- H.E.S.S. algorithm implemented (Chaudhuri-Kundu and 2D log-likelihood) with preliminary tests
- “muonmaster” branch no longer exists → has been merged into “master” branch
- For implementing more methods, **please create a new branch for each method**
  - Easier to keep track of and review changes
  - Cleaner approach, smaller steps, more frequent merging
- Muon\_reconstruction example has been adopted into a tool (MuonDisplayTool)

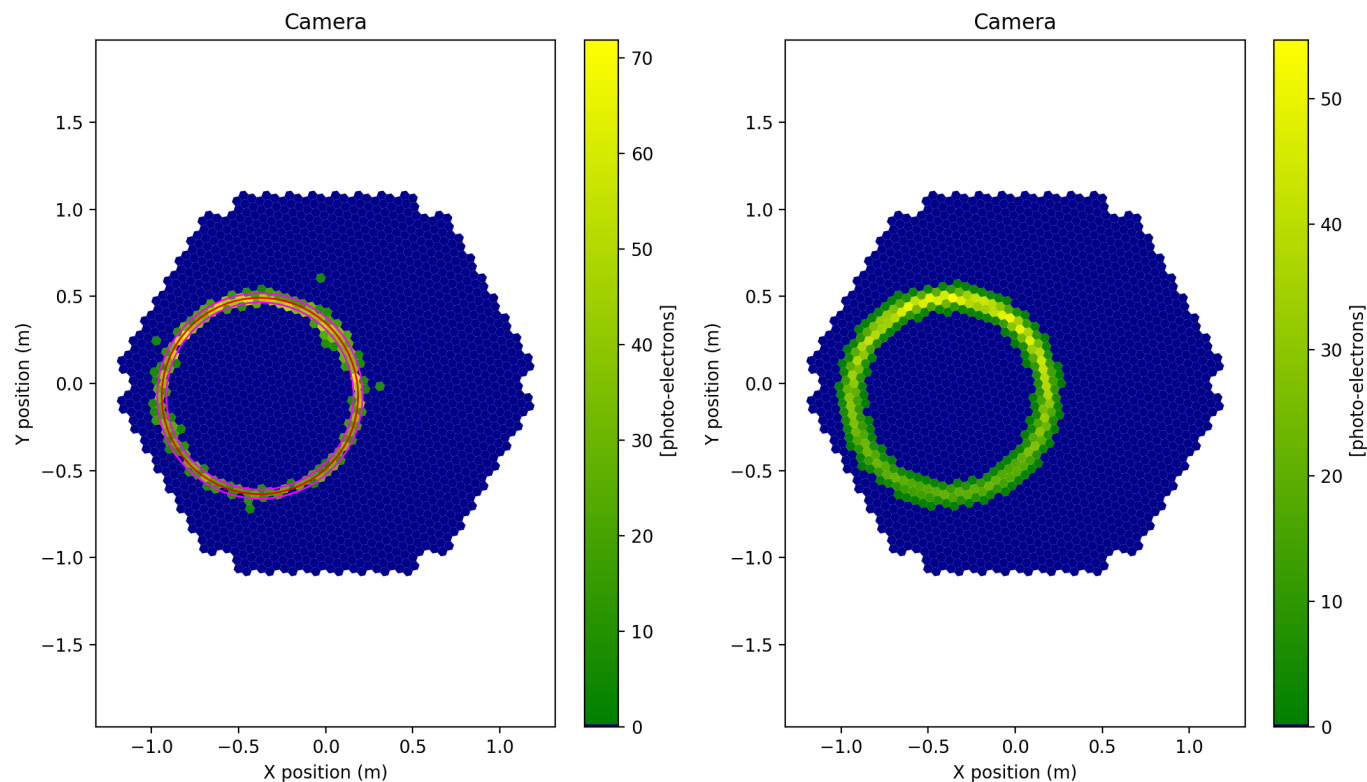


# List of relevant files in ctapipe

- `examples/muon_reconstruction.py`
- `image/muon/`
  - `muon_reco_functions.py`
  - `muon_ring_finder.py`
  - `muon_integrator.py`
  - `muon_diagnostic_plots.py`
- & generic classes: `intensity_fitter.py` & `ring_fitter.py`
- MC muon simulation test files `mst-fc` and `sst-dc` at 100% efficiency in `ctapipe-extra` repository  
(can add/distribute more xSTs / efficiencies if there is interest → would need to as Karl Kosack)

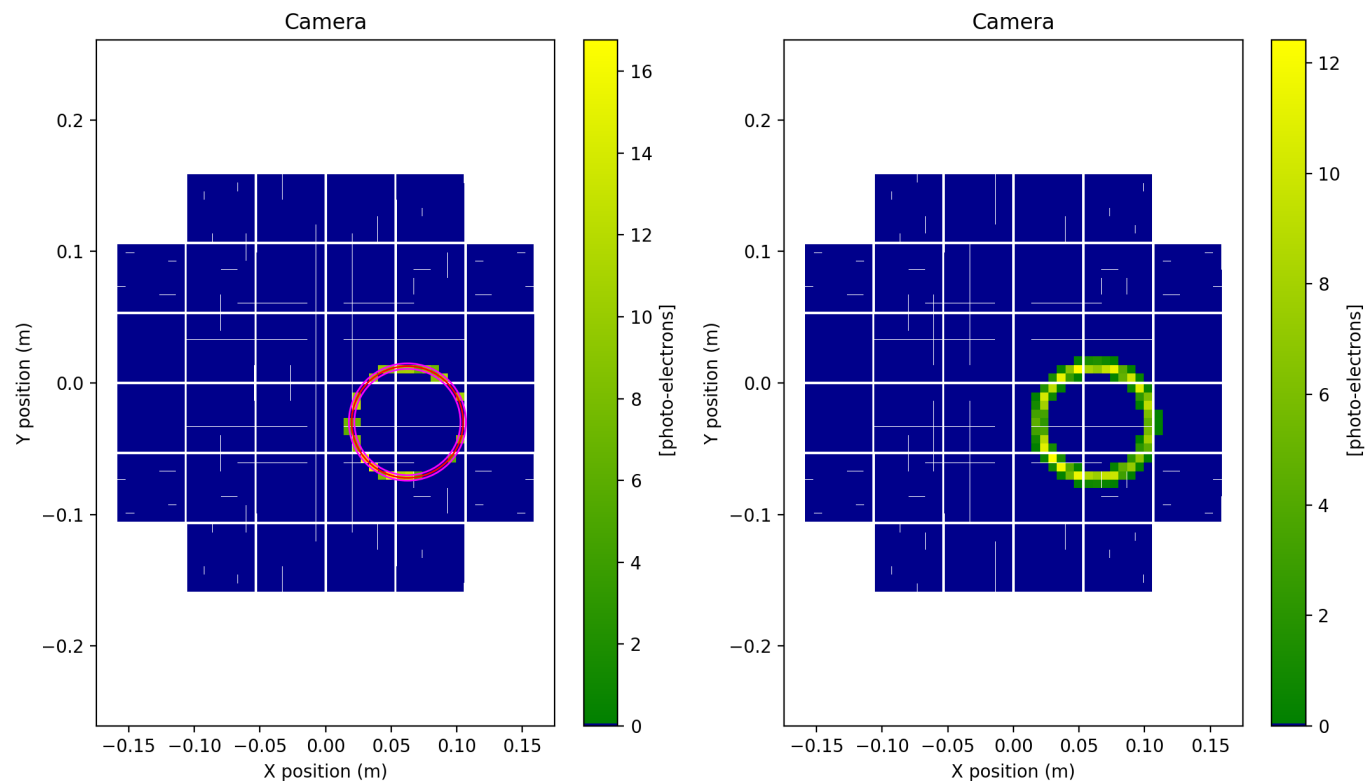
# Muon Image Fitting

- Circle and width fit, intensity expectation



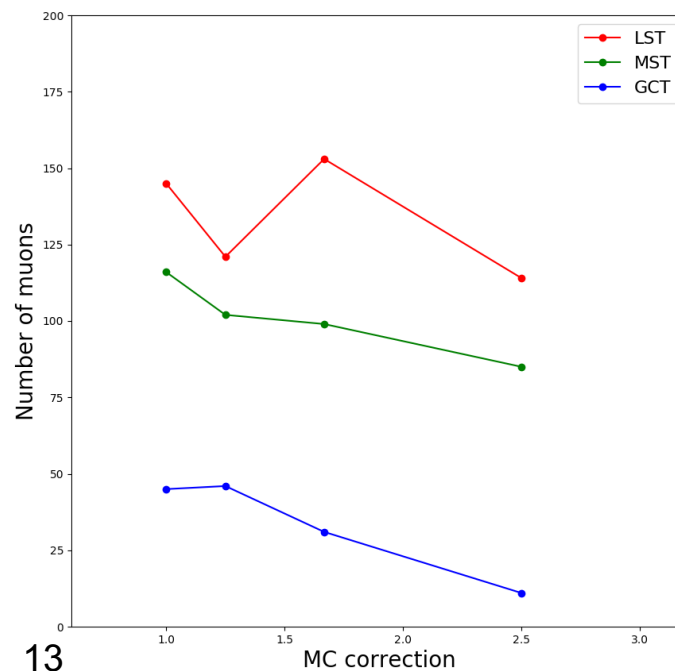
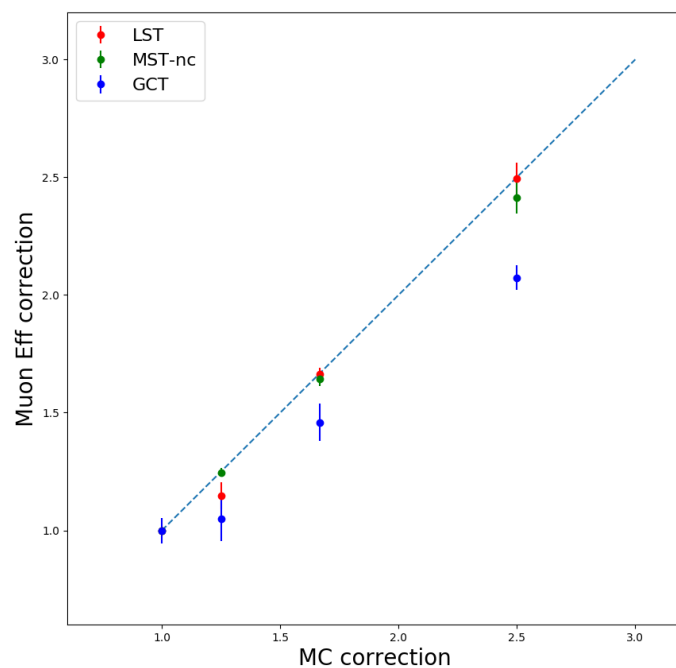
# Muon Image Fitting

- Circle and width fit, intensity expectation



# Figures of Merit – so far

- Gaussian-like efficiency distribution?
- Linearity of correction factor?
- Number of muons found?



# Other Figures of Merit

- Independence of muon ring width / impact parameter from the muon optical efficiency
  - can be tested
- Robust performance against broken pixels
  - Needs dedicated study / extra MC
- Agreement between single muons and hadronic muons (e.g. proton simulations)
  - Can be tested with hadronic MC
- Independence of muon efficiency from observation position / environment
  - Probably needs data

# Muon Calibration – to do

- Write more unit tests in ctapipe  
(throw warnings if code doesn't work as expected)
- Implement other algorithms
- Check and improve figures of merit
- Implement other applications of muons
  - Psf, camera flat-fielding, mirror calibration,
- Investigate other muon issues
  - Bias due to hadron / wavelength/ broken pixels...



# Muon Calibration Paper - Suggestion

- For a technical study, compare the known methods of muon reconstruction and draw conclusions on “best” approach, e.g.
  - Most muons usable for calibration
  - Linearity with optical efficiency
  - Gaussianity of distribution
  - Dual-mirror telescopes performance
- To compare in as unbiased an approach as possible:
  - Implement all algorithms for comparison in ctapipe
  - Use same simulations for testing all algorithms
  - Follow standard input/output/storage class format





# Thank you for your attention

Any Questions?

