

# Modeling Particles in the ISM

Including the ISM's Influence on the Spatial Distribution of Protons

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# **Gamma-Ray Production**





# Modeling particles in the ISM

#### Sabrina Einecke @ CTA-Oz Nov. 2021 https://indico.cta-observatory.org/event/3712/contributions/31509/





9°00' 8°30' 00' 7°30'

Galactic Longitude

00' 7°30'

Galactic Longitude

00

Modeling Particles in the ISM | S. Einecke

Conclusion

Future additions:

Neutrinos

Future plans:

partISM:

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# **Motivation**



### **Current state:**

Diffusion coefficient (and therefore diffusion lengths)

the same for the whole map

### **Reality:**

Diffusion coefficient position dependent



**Diffusion Coefficient** 

Can be estimated by number density of ISM (Crutcher)

### **Expectation:**

- Diffusion length decreases with increasing magnetic field
- Higher proton/gamma-ray density in areas of higher magnetic fields

# Regions with different number densities

Accelerator in the centre

### **Comparison of models - Gas map**





# **Comparison of models - Proton Distribution**





# **Comparison of models - Gamma-ray Flux**





# Conclusion



### Achievments:

- Implementation of position-dependent diffusion lengths in partISM
- Obtaining a position-dependent proton distribution

### **Next steps:**

- Use a numerical solution of the transport equation for a comparison model
- Validate my analytical model on the results of the numerical solution

### **Future plans:**

- General modeling of all SNRs
- Detailed modeling of few regions with PeVatron candidate SNRs